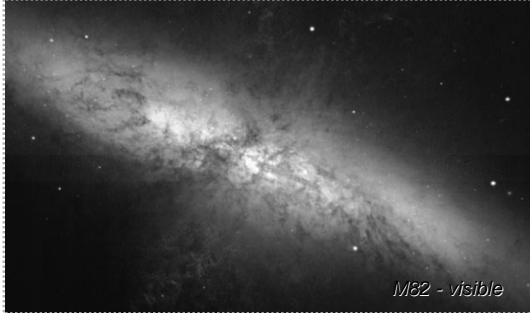


ASTR 1120: Stars & Galaxies



Prof. Juri Toomre TA: Ben Brown/Amy Bender
Lecture 36 Mon 11 Apr 05
zeus.colorado.edu/astr1120-toomre

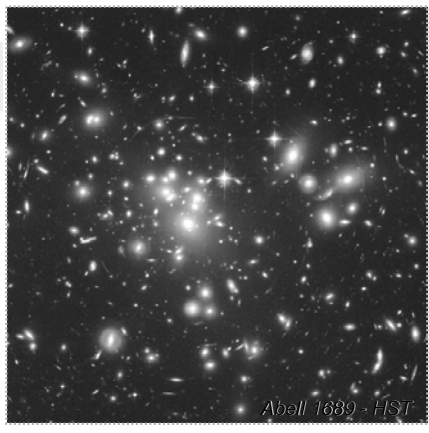
Today's "Ride to the Wild Side"

- Today discuss *active galaxies* – quasars, *starburst galaxies*, *radio galaxies*
- All from *Chap 21: Galaxy Evolution* – start overview reading *Chap 22: Dark Matter* for Wed lecture
- Third Mid-Term Exam this Friday 15 April
- Review Set 3 available, Ben Brown runs evening review this Wed 7-9pm
- Homework Set 9 Planet Finder closes Wed

Rich galaxy cluster

Galaxies are not lonely
-- many "interactions" most likely!

also curious arcs of light?



Reading clicker – the boss galaxy

- Which of the following is NOT a feature of a central dominant galaxy ?

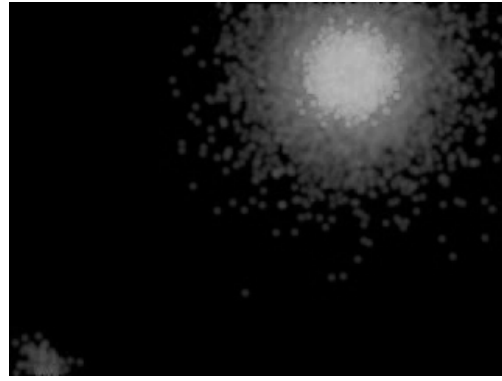
A.

- A. They are often spiral galaxies
- B. They are found in clusters of galaxies
- C. They often have multiple galactic nuclei near their centers
- D. They are thought to form by the merger of several smaller galaxies

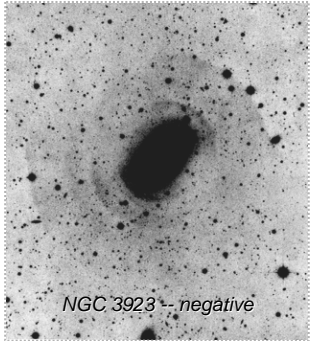
Simulation: tidally interacting galaxies Barnes



Another pair -- galaxies tumble together Barnes



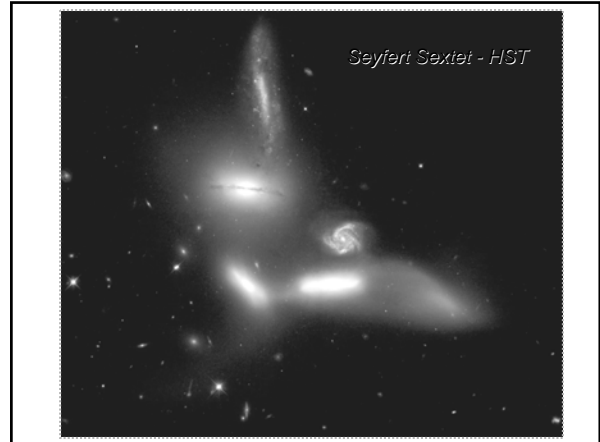
Signs of having collided:
elliptical galaxy with shells



NGC 3923 -- negative



"Polar ring" galaxy

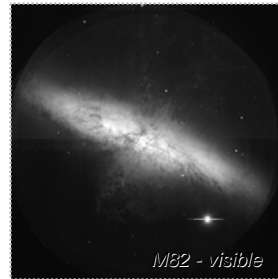


Seyfert Sextet - HST

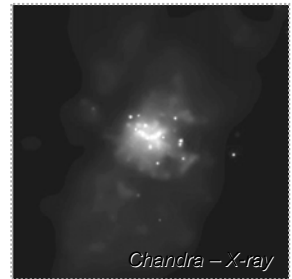
Messages from galaxy interactions

1. In dense clusters, galaxy collisions (grazing or even head-on) must have been common
2. With successive passages, spiral galaxies can tumble together to form a big elliptical
3. Vastly increased star birth from shocking the gas and dust (star burst galaxies)
4. Start rapid feeding of supermassive black hole lurking at center of most galaxies (quasars)

Starburst Galaxies



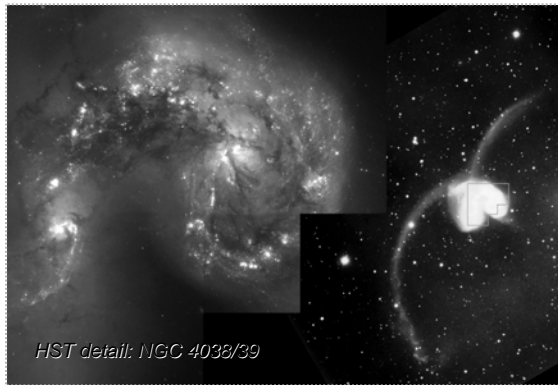
M82 - visible



Chandra - X-ray

- Milky Way forms about 1 new star per year
- Starburst galaxies form 100's of stars per year

Vigorous star birth – "The Antennae"

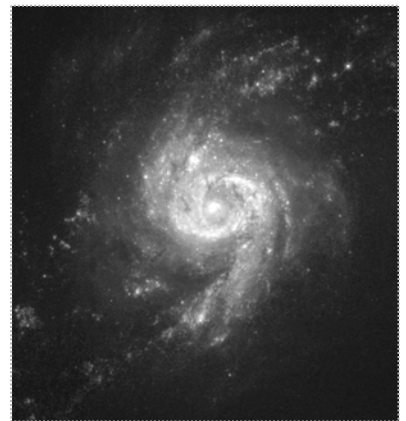


HST detail: NGC 4038/39

Starburst galaxy in fine detail

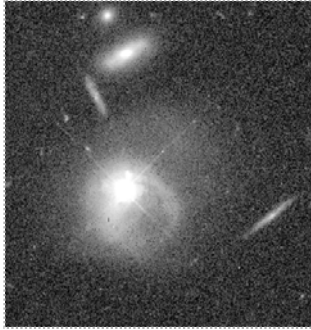
NGC 3310 - HST

Big open two-sided spiral structure
--> tidal interaction



Quasars

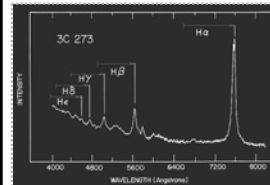
- Quasi-stellar Radio Source (QSOs)
- Nuclei so bright that the rest of the galaxy is not easily seen
- First discovered as radio sources - then found to have high redshifts! (far, far away?)



QUASARS – what are they?

“3C 273”
redshift $Z = 0.16$

1000 times brighter than Milky Way!



WHAT IS A QUASAR? (QSO)

QUASI-STAR RADIO SOURCE OBJECT

1960: 3C 49 FIRST QSO DISCOVERED (5th CAMBRIDGE RADIO CATALOG) $z_{obs} = 0.262$

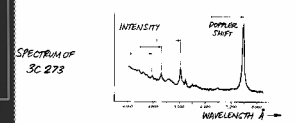
1963: 3C 273 MARSHEN SCHMIDT (CALTECH) SEARCHED OUR PORTALS OF SPACE— VERY LARGE RED SHIFTS IN HYPOTHESIS! EMISSION LINES— BROAD & DIMINISHED POWER SHIFT = 15% C (MORSE LICH) ($v = 0.15c$; 3 B LY AWAY)

QSO:

TOTAL OPTICAL LUMINOSITY (IF AT COSMOLOGICAL DISTANCES) $\sim 10^{13} L_{\odot}$

NORMAL GALAXY $\sim 10^{10} L_{\odot}$

ABOUT 10% QSO'S ARE RADIO SOURCES



Cosmological (Big) Redshifts (from expansion of universe)

Alternative definition of redshift :

$Z = \text{redshift}$
= change in wavelength / “normal” wavelength

$1 + Z = \frac{\text{observed wavelength}}{\text{“normal” wavelength}}$

redshifts always have $Z > 0$
(redder light has larger wavelengths)

Quasar spectra

QUASAR SPECTRA

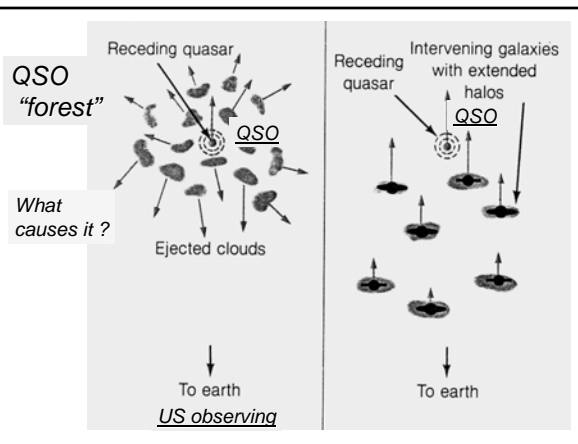
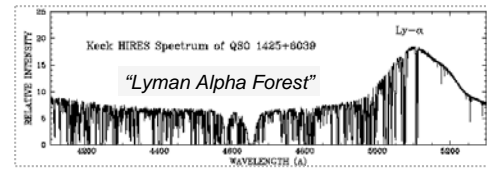
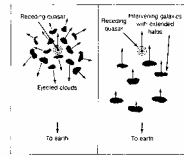
• REDSHIFTS DETERMINED FROM SYSTEMIC EMISSION LINES OF HIGHLY IONIZED GAS

QSO: CONTACT BLUE OBJECT, SOME WITH RADIO & X-RAY EMISSION

• BUT ALSO WHOLE “FOREST” OF ABSORPTION LINES AT VARIOUS SMALLER REDSHIFTS

EMERGES:
• ABSORPTION IN CLOUDS EJECTED FROM QUASAR OR (MORE LIKELY)
• FROM EXTENDED HALOS OF INTERVENING GALAXIES

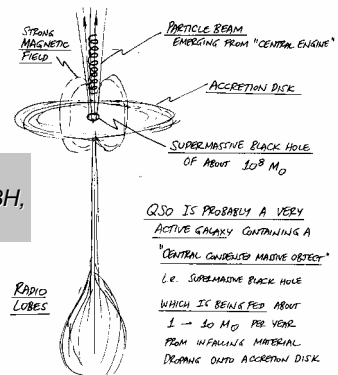
Redshifts from emission lines + Many absorption lines (forest)



Model for “active galaxies”

MODEL OF RADIO GALAXIES, QUASARS, SEYFERTS

Accretion disk, supermassive BH, beams on axis



"Central Engine" -- artist's conception

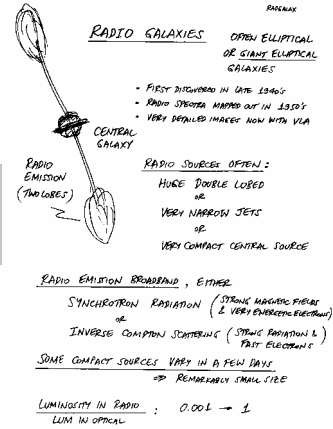
- Accretion disk around super-massive black hole
- Disk itself may or may not be obscured by dust
- If bright nucleus is visible, looks like a quasar, if not, then a radio galaxy



Radio galaxies

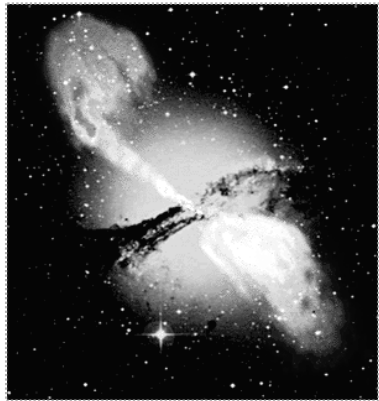
Central elliptical galaxy, huge lobes of emission, compact central source

Synchrotron radiation

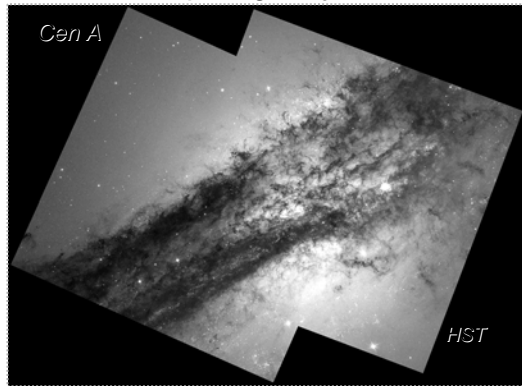


Prototypical "radio galaxy"

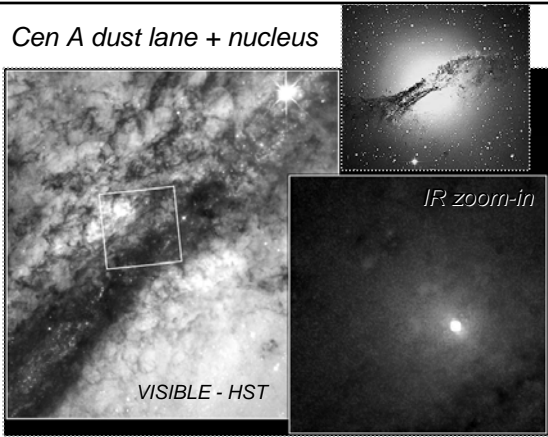
Giant elliptical galaxy NGC 5128 with dust lane (from spiral galaxy?) + Centaurus A radio source (color lobes)



Remains of spiral galaxy as dust lane ?



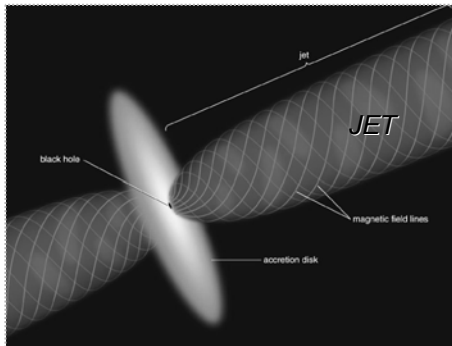
Cen A dust lane + nucleus



Clicker – what makes the light?

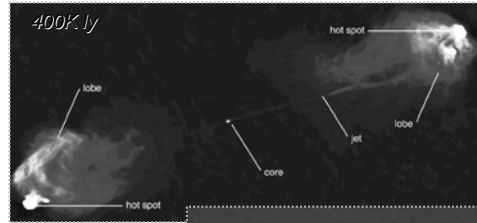
- What is most likely source of light from bright nuclei (radio, visible, x-rays) in active galaxies?
- D.**
- A. Thermal radiation from a massive star cluster
 - B. Emission lines from hot gas
 - C. 21 cm from hydrogen
 - D. Synchrotron radiation from a black hole

Synchrotron radiation from particles moving outward

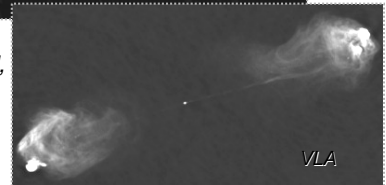


Spinning accretion disk drags along magnetic fields

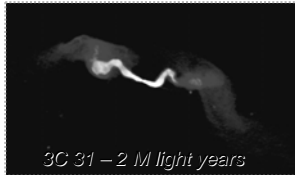
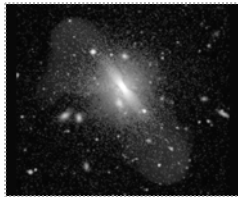
Cygnus A radio jets



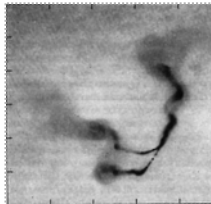
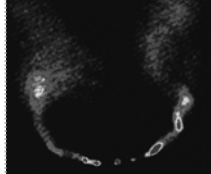
Jet as fine thread,
big lobes at end,
central hot spot



Radio tails: many shapes



NGC 1265 - 100K ly



Typical properties
of "active galaxies"

synchrotron
emission !

source very
small in size

PROPERTIES OF "ACTIVE GALAXIES"

RADIO GALAXIES, SEYFERT GALAXIES,
BL LACERTAE OBJECTS, QUASARS "SOME HAVE MORE
THAN OTHERS!"

1. HIGH LUMINOSITY
= MUCH MORE LUMINOUS THAN NORMAL GALAXIES
2. NON-THERMAL EMISSION
 - EXCESS RADIATION IN UV, IR, RADIO, X-RAY
 - INCLUDING SYNCHROTRON EMISSION FROM RELATIVISTIC ELECTRONS SPIRALING IN MAGNETIC FIELDS
3. SMALL, COMPACT SIZE OF INTENSE EMISSION
= NUCLEUS VERY BRIGHT COMPARED TO REST OF GALAXY
4. RAPIDLY VARYING EMISSION
• SOURCE MAY BE A FEW LIGHT HOURS OR DAYS IN SIZE
5. EXPLOSIVE FEATURES
• JETLIKE EXTENSIONS, FILAMENTS
6. GRAVITATIONAL DISTURBANCES
 - VERY HIGH INTERNAL VELOCITIES PRODUCED FROM BROAD SPECTRAL EMISSION LINES
 - PECULIAR OPTICAL APPEARANCE
7. LARGE REDSHIFTS
= IMPLYING HIGH RECESSIONAL VELOCITIES, VERY LARGE DISTANCES