

Today's Topics

- Today begin looking at Chap 21 Galaxy Evolution, especially at active galaxies
- Particularly important is that <u>galaxies can</u> <u>experience collisions</u>: "interacting galaxies"
- Read 21.4 Starburst galaxies and 21.5 Quasars and active galactic nuclei in detail for Monday
- <u>Third Mid-Term Exam</u> next <u>Fri 15 April</u> (tax day!) <u>Review Sheet 3</u> available today; <u>Homework 8</u> returned graded + answers

Measuring big distances to galaxies

"STANDARD CANDLES" -- important ones in `distance ladder', or `chain'

REMINDER

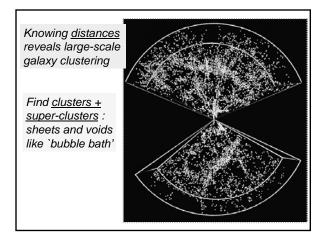
- 1. Main-sequence fitting
- 2. Cepheid variables
- 3. Tully-Fisher relation
- 4. White dwarf supernovae

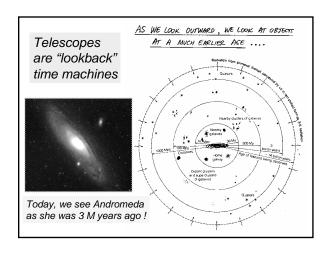
Brightness ~ Luminosity / (Distance)²

DISTANCE ESTIMATE 5 Use <u>Hubble's Law</u> itself to estimate vast distances D

- Measure velocity, then: $D = v / H_0$
- Example: using H_o = 70 km/sec/Mpc, and finding that v = 700 km/sec

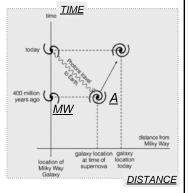
D = 700 km/sec / 70 km/sec/Mpc = 10 Mpc= 32 million light years





Lookback time (in expanding universe)

- Say it takes 400
 million years for light
 to get from galaxy A
 to us in Milky Way
- Yet during travel in <u>spacetime</u>, both <u>A</u> and <u>MW</u> have changed positions by expansion
- Thus "distance" is a fuzzy concept – LOOKBACK TIME is better

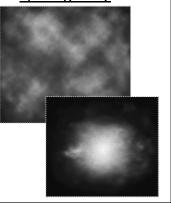


Clicker on reading ahead

- D.
- What do we mean by a protogalactic cloud?
- A. It is a cloud-like halo that surrounds the disks of spiral galaxies
- B. It was a term used historically to refer to any galaxy
- C. It is a cloud of hydrogen gas that we detect by looking at light from quasars
- D. It is a cloud of matter that contracts to become a galaxy

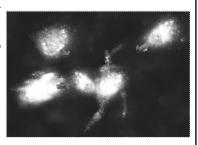
Making of a spiral galaxy

- Start with a fairly uniform cloud of hydrogen
- Gravitational collapse forms protogalactic clouds
- First stars are born in this spheroid (such stars are billions of years old → "fossil record")



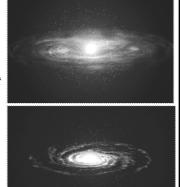
Small variant in spiral making ...

- Several smaller protogalactic clouds may have merged to form a single large galaxy
- May explain slight variations in stellar ages in the MW



Forming a <u>disk</u> <u>with spiral</u>

- As more material collapses, angular momentum spins it into a disk
- Stars now formed in <u>dense spiral</u> <u>arms</u> – disk stars are younger!



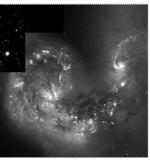
Making ellipticals

- Higher density: much faster star formation uses up all the gas
- Nothing left to make a disk
- Now we see sphere of old stars

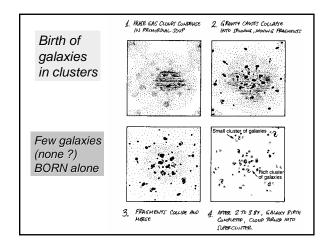




- Spiral galaxy collisions destroy disks, leave behind elliptical
- Burst of star formation uses up all the gas
- · Leftovers: train wreck
- Ellipticals more common in dense galaxy clusters

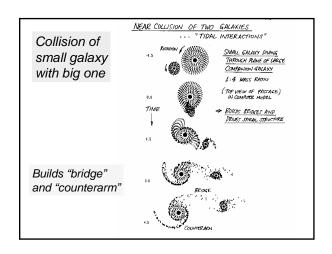


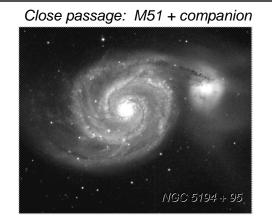
NGC 4038/39 Antennae

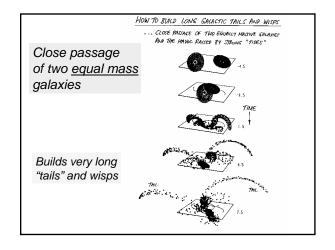


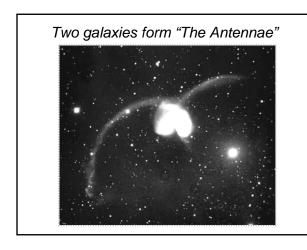
Clicker: galaxy collisions

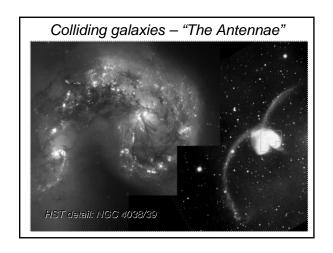
- Why are <u>collisions between galaxies</u> more likely than between stars within a galaxy?
- A. Galaxies are much larger than stars
- B. Galaxies travel through space much faster than stars
- C. Relative to their sizes, galaxies are closer together than stars
- D. Galaxies have higher redshifts than stars

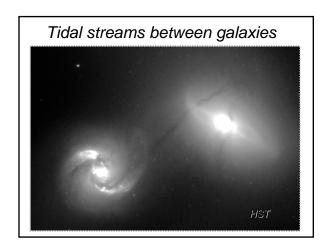


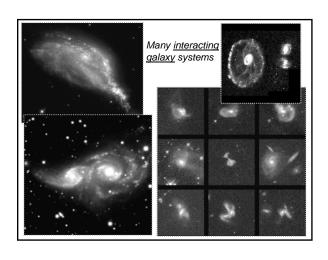


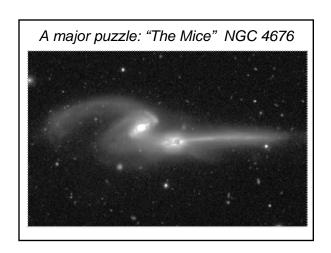


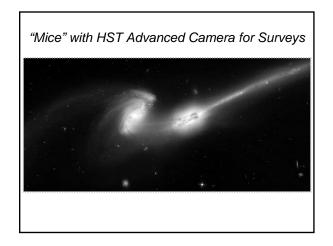


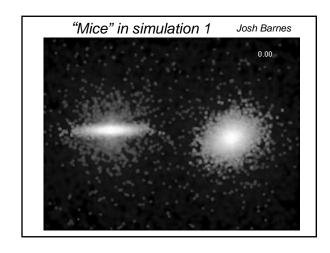


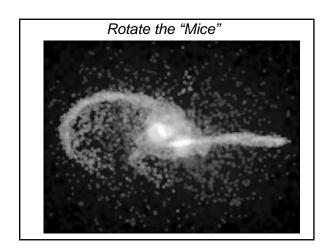


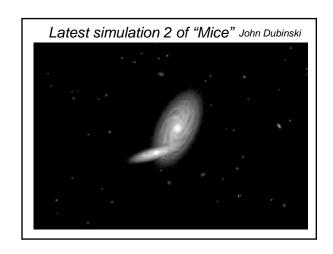


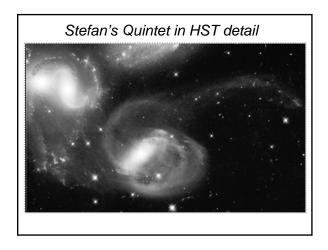




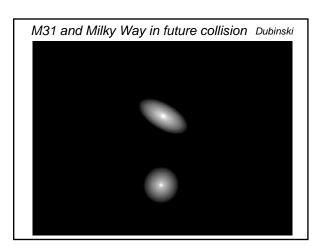














Quasars

- Quasi-Stellar Radio Source (QSO) – arise from early galaxy collisions feeding BH?
- Nuclei so bright that the rest of the galaxy is not easily seen
- First discovered as radio sources then found to have <u>very</u> high redshifts!

