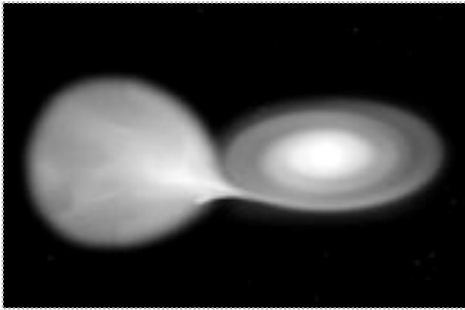


ASTR 1120: Stars & Galaxies



Prof. Juri Toomre TA: Ben Brown
Lecture 24 Mon 7 Mar 05
zeus.colorado.edu/astr1120-toomre

More Truly Bizarre Events

- *White dwarf supernovae* from mass transfer in binary system, but also repeated novae
- *Hot accretion disks* formed around object (white dwarf, neutron star, black hole) on receiving end of mass transfer
- Why more pulsars than supernova remnants (SNR)
- *Black holes*, their general properties, and their "care and feeding"
- *Second Mid-term Exam* this Friday, review session Wed 7pm, Review Set 2 available
- *Homework Set 6* due today

Review – what is left behind?

- After a *massive-star* supernova explosion, the remains of the stellar core _____ .
- B.**
- A. will always be a neutron star
 - B. may be either neutron star or black hole
 - C. will always be a black hole
 - D. may be either a white dwarf, neutron star, or black hole

FINAL FATE OF MASSIVE STARS

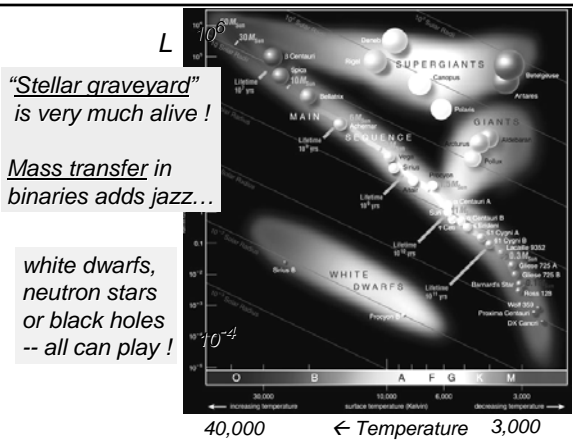
Final fate of massive stars

.... RATHER UNCERTAIN IN DETAILS THAT DEPEND ON MASS LOSS BY STELLAR WINDS
INFLUENCED BY { INITIAL MASS M
IRON CORE MASS M_{core} }

CURRENT THINKING (THEORY/COMPUTER MODELLING)....

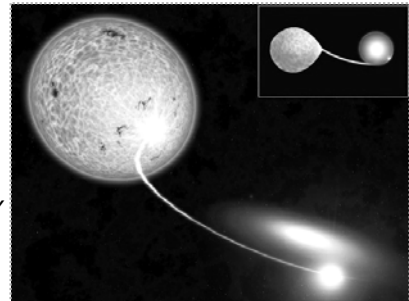
1. IF HEAVYWEIGHT $M \leq (20-30) M_{\odot}$:
EVOLUTION YIELDS $1.4 M_{\odot} < M_{core} < (2-3) M_{\odot}$
⇒ CORE COLLAPSE & CORE BOUNCE
⇒ **Neutron star**
SUPERNOVA (TYPE II)
⇒ NEUTRON STAR

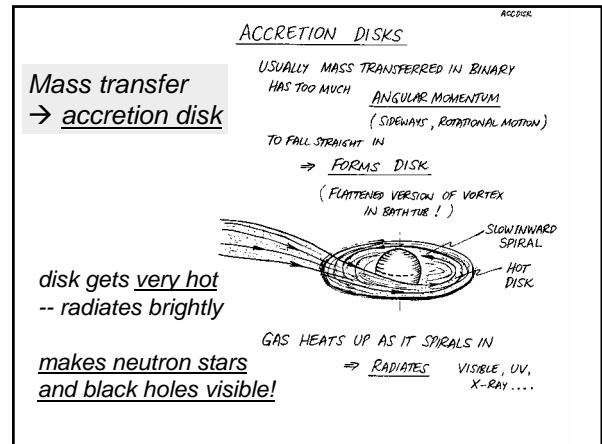
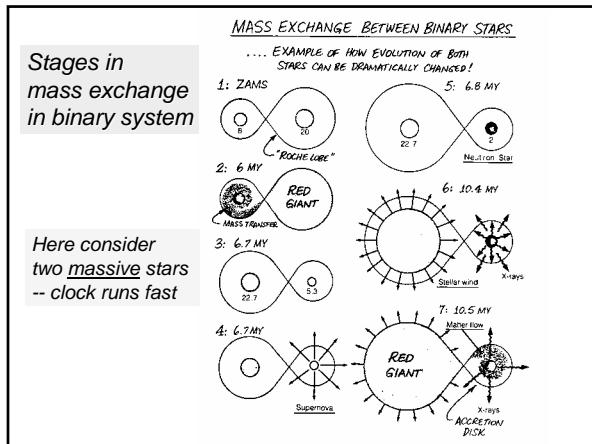
2. IF TRULY MASSIVE $M \geq (20-30) M_{\odot}$:
THEN PROBABLY GET $M_{core} > 3 M_{\odot}$
⇒ COLLAPSE AND SUPERNOVA EXPLOSION
⇒ **BLACK HOLE**



"Compact Companions" in Binary Systems

- Mass transfer from red giant companion spirals onto accretion disk
- Inner parts become VERY hot -- glow in UV, X-rays





White Dwarfs in Binary Systems

- Again, mass transfer from red giant companion spirals onto an accretion disk
- But too much mass can take white dwarf over the edge!

WHITE DWARF PYROTECHNICS

WD ALONE ... BORING $e^2 e^2$ BUT IN CLOSE BINARY, WOW!

TYPE WD: CARBON & OXYGEN COMPOSITION: H & He

MASS TRANSFER (ACCRETION) DUMPS H & He ONTO WD SURFACE UNTIL ... THREE POSSIBILITIES:

- LOCALIZED NUCLEAR FLASH BURNING ON SURFACE (INTERMITTENT) ⇒ "CATACLYSMIC VARIABLE STAR"
- ENOUGH "FUEL" PILES UP TO IGNITE INTENSE CNO CYCLE ... EXPLOSION BLOWS OFF (EJECTS!) OUTER LAYER ⇒ NOVA ("NEW", OR "GUEST" STAR)

BRIGHTENS TO $\sim 10^5 L_{\odot}$ FOR FEW WEEKS, THEN FADES (CAN RECUR)

NOVA

NOVA HERCULIS (1935)

DARKEST IN MARCH, FIRED AWAY IN MAY, BLOWN OFF SHELL [IN 1972]

Nova

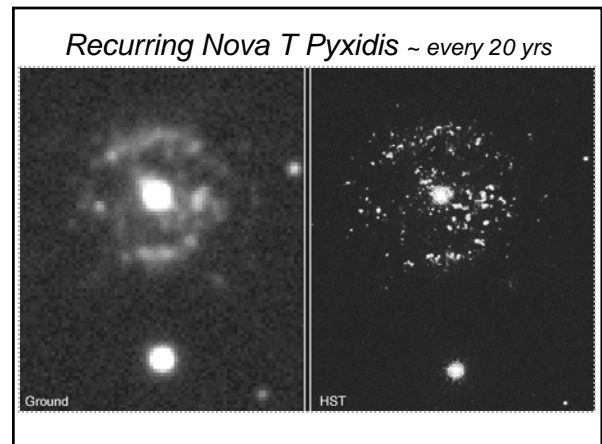
- Accretion of gas onto white dwarf can lead to H fusion on surface
- Star becomes much brighter → nova (may blow off shell)

white dwarf, companion star

Hydrogen-rich gas spills into an accretion disk and forms a shell of hydrogen on the white dwarf.

A nova occurs when the shell becomes hot enough for a burst of hydrogen fusion.

Nova Cygni 1992+2



White Dwarf SUPERNOVA

3: If exceed $1.4 M_{\text{SUN}}$

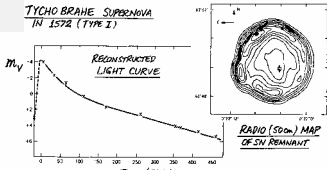
Collapse of WD, explosive fusion burning of "carbon star" – all gone!

Brightest SN: superb beacons for measuring distances

WHITE DWARF SURPRISES...

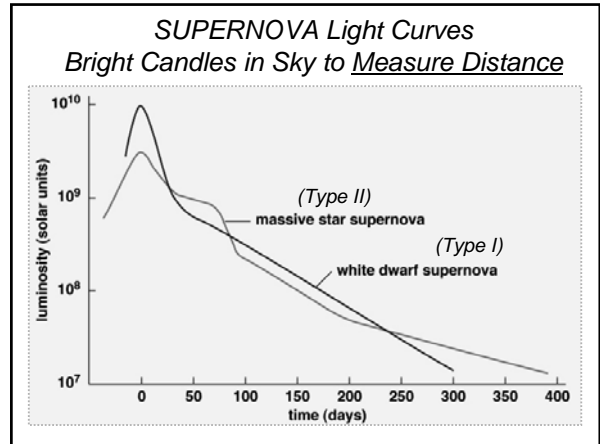
3. IF WD CLOSE TO $1.4 M_{\odot}$ LIMIT, ACCRETED MASS MAY TAKE IT "OVER THE EDGE"
 ⇒ STAR BEGINS TO COLLAPSE, INTERIOR HEATS UP, EXPLOSIVE NUCLEAR BURNING OF CARBON!...
 ENTIRE STAR BLOWS APART!
 ⇒ SUPERNOVA (TYPE I, NO H LINES)
 BRIGHTENS TO $10^9 L_{\odot}$ (BRIGHTER OF ALL!) FOR FEW WEEKS
 NOTHING LEFT BUT EXPANDING SHELL (NO NEUTRON STAR)

TYCHO BRAHE SUPERNOVA IN 1572 (TYPE I)

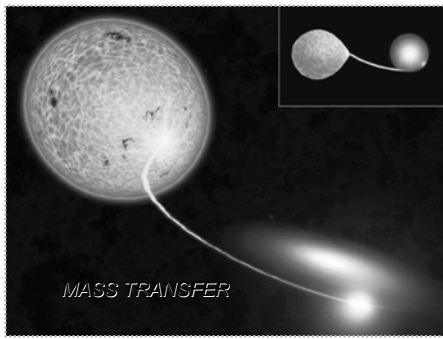


RECONSTRUCTED LIGHT CURVE

RADIO (50cm) MAP OF SN REMNANT



Since white dwarfs in evolving binary systems come "alive" – what about neutron stars?



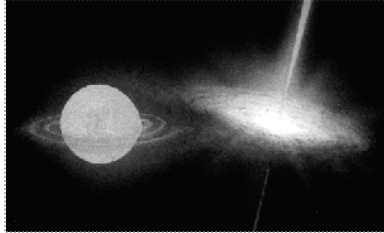
MASS TRANSFER

Binary WD:
Hot accretion disks, novae, supernovae

Neutron star:
Radiation with more vigor, no SN

Neutron Stars in Binary Systems

If white dwarfs can do it, so can neutron stars!



- Mass transfer builds very hot accretion disk around neutron star:
 - intense x-ray emission (continuously)
 - explosive helium burning (in bursts) on disk

"X-Ray Bursters"

X-RAY BURSTERS: ACCRETION FLOW FROM BINARY COMPANION ONTO NEUTRON STAR ⇒ EXPLOSIVE HELIUM BURNING ⇒ X-RAY BURSTS

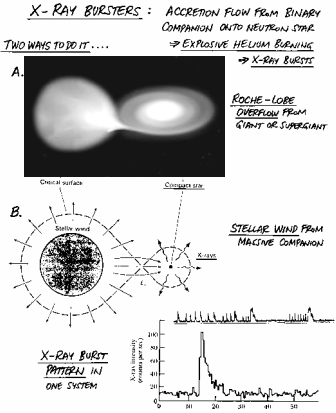
TWO WAYS TO DO IT...

A. **ROCHE-LOBE OVERFLOW FROM STAFF OR SUPERGIANT**

B. **STELLAR WIND FROM MASSIVE COMPANION**

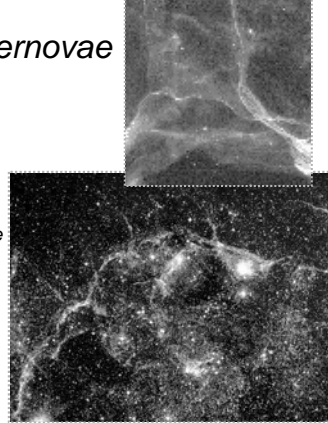
Accretion flow onto neutron star (overflow or wind) → helium fusion flash burning

X-RAY BURST PATTERN IN ONE SYSTEM



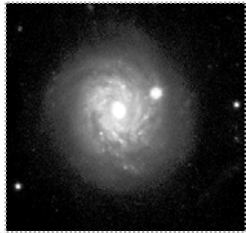
Observing Supernovae

- About 1 per century per galaxy (none in Milky Way since 1604) ☹
- Bright explosion visible for weeks/months - some visible in daytime!
- Remnant visible for 10,000+ years as huge bubbles and "veils" – longer in radio



SUPERNOVAE in Other Galaxies

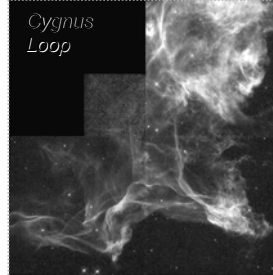
- Bright enough to be seen as sudden, bright point in other galaxies
- Many astronomers monitor nearby galaxies nightly to catch them
- 1 per 100 years per galaxy means that if you monitor 100 galaxies, see ~ 1 SN per year)



SUPERNOVA REMNANTS (SNR)

SNR "numbers game"

- ... EXPANDING DEBRIS OF EXPLOSION
- DO NOT LAST VERY LONG BEFORE DISSIPATING
- FOR ABOUT 1000 YEARS, SEEN IN X-RAY, VISIBLE RADIO
 - AFTER ~ 10,000 YRS, SHOCK WAVE SIGNIFICANTLY WEAKENED, BUT GAS HAS COOLED
 - RADIO EMISSION AGAIN
 - PROBABLY INVISIBLE AFTER ~ MILLION YEARS



BUT... SUPERHELIX CAUSED BY SUCCESSIONAL SN IN OR ASSOCIATION MAY LAST MUCH LONGER!

THE NUMBERS GAME:

- ~ 100 SNR DETECTED IN OUR GALAXY (MOST IN RADIO)
- 5 HAVE IDENTIFIABLE PULSARS
- ~ 50 SNR IN LMC
- ~ 10 SNR IN SMC
- BUT ~ 400 PULSARS FOUND IN OUR GALAXY, 1 IN LMC
- ANSWER: PULSARS WERE BORN (~ 4 MY), PROBE MOTION OF PULSAR MAY CHECK IF AWAY FROM SNR

Review of "COMPACT OBJECTS"

all charming but bizarre!

"COMPACT OBJECTS"

WHITE DWARFS

- RADIUS: ~ 5,000 Km (~ EARTH)
- MASS: < 1.4 M_{\odot} (CHANDRASEKHAR LIMIT)
- DENSITY: 10^6 X WATER
- OBSERVED: INDIVIDUAL, AND IN BINARIES
- ⇒ NOVAE, CATAclysmic VARIABLES

NEUTRON STARS

- RADIUS: ~ 10 Km (BOUNCE → NAVY?)
- MASS: < 2-3 M_{\odot}
- DENSITY: 10^{15} X WATER (NUCLEAR DENSITY!)
- OBSERVED: PULSARS, X-RAY BINARIES

BLACK HOLES

- RADIUS: "HORIZON" DEPENDS ON MASS
- MASS: ANY
- DENSITY: DEPENDS ON MASS
- OBSERVED: ? SEVERAL CANDIDATES

Clicker Review – White Dwarfs

- A white dwarf (WD) is gaining mass because of accretion in a binary system. What happens if its mass reaches the 1.4 solar mass limit? **C.**
- A. WD will collapse to become a black hole
- B. WD will collapse, becoming a neutron star
- C. WD will explode completely as a WD supernova
- D. WD will undergo a nova explosion

Reading Needed

- Complete detailed reading of 18.4 Black Holes
- Overview read S3 Spacetime and Gravity, and more carefully S3.4 New View of Gravity

Black Holes – sort of courtesy of Albert

- Einstein's (1911) General Theory of Relativity: gravity is really the warping of spacetime around an object with much mass
- Light travels in "straight lines" – and its bending comes from spacetime being curved by gravity

