

#### Our Schedule

- Mid-Term Exam 3 essay grading not yet finished (from "L" onward) – will post details
- Homework #13 due today (flexible)
- Please turn in extra-credit Observatory Night project by next Tues Apr 28
- Please do course evaluation (FCQ) online for course + recitation this week
- Focus on 22.2 Evidence for Big Bang and on 22.3 Big Bang and Inflation
- Complete detailed read Chap 23: Dark Matter, Dark Energy, Fate of Universe

## Cosmology topics and issues

- · Look at models for our universe
- Cosmic microwave background (CMB) and all its implication
- Ideas of "dark energy" arising from:
- White-dwarf supernova data
- CMB mapping
- Imply "accelerating universe"

### REMINDER

<u>Cosmological (Big) Redshifts</u> (from expansion of universe)

Alternative definition of redshift:

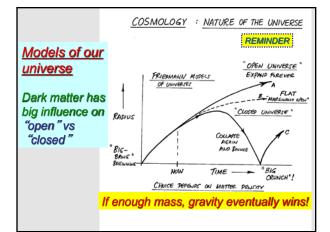
Z = redshift

= change in wavelength/ "normal" wavelength

1 + Z =

observed wavelength / "normal" wavelength

redshifts always have Z > 0 (redder light has larger wavelengths)



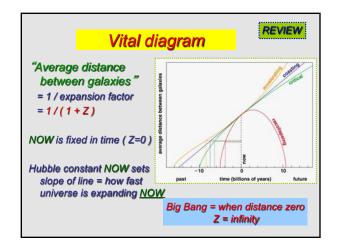
### REMINDER

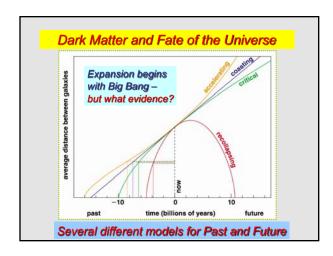
## Predictions of General Relativity Theory (GRT)

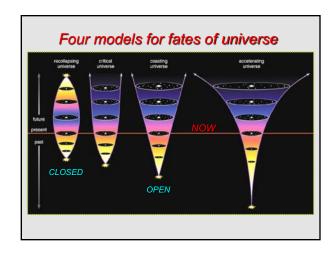
- <u>Einstein</u> in 1917 realized GRT (1915) predicted universes in motion, but preferred `steady state' – added `cosmological constant' (CC) as repulsive force in space-time to counteract attractive force of gravity
- Willem de Sitter (A, Dutch, 1917) solves GRT equations with no CC and low density of matter: showed universe must expand
- Alexander Friedmann (M, Russian, 1920) solves GRT with no CC but any density of matter: universes can expand forever, or collapse again, depending on mean matter density

# More on ... Predictions of GRT

- Georges Lemaitre (P, Belgian, 1927) rediscovers
   Friedmann solutions, told Hubble (observing
   redshifts since 1924) that cosmic expansion
   suggests more distant galaxies should have greater
   redshifts (Hubble publishes V = H<sub>o</sub> d in 1929)
- <u>Einstein visited Hubble in 1932</u>, said CC "biggest blunder"





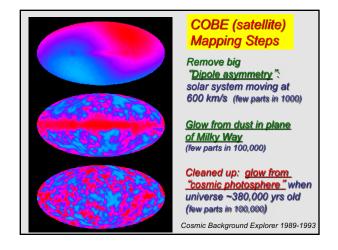


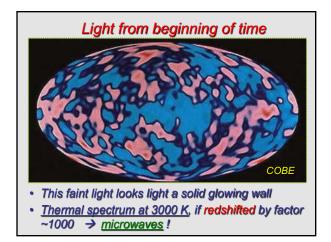
Questions or Comments

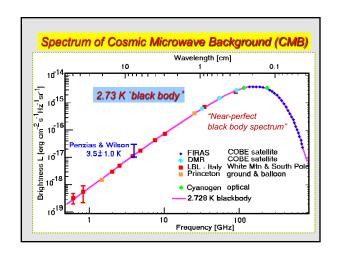


### CMB (Accidental) Detection Story

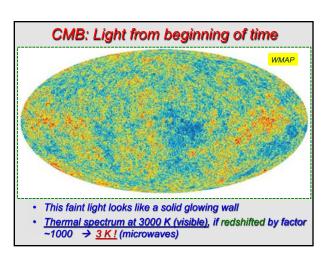
- George Gamow, Robert Dicke and Jim Peebles are some players in predicting (1946-1960s) that a remnant radiation signal (microwave background temperature) should survive from "Big Bang" beginning of universe
- Spectrum "temperature" estimates ranged from 50K to 20K or less
- <u>Robert Dicke</u> at Princeton in 1964 was building a horn with his earlier WWII design (Dicke radiometer) to look for background microwave radiation
- <u>Arno Penzias and Robert Wilson</u> at nearly same time used big horn antenna at Bell Labs (with cooled Dicke radiometer) to start radio mapping of Milky Way
- Their "background noise" at 4000 MHz (7.35 cm) was inexplicable – Bernie Burke told them to talk to Dicke!

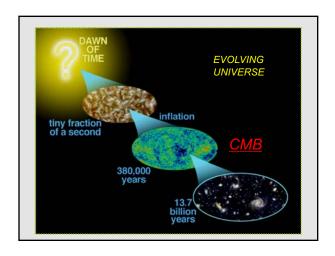






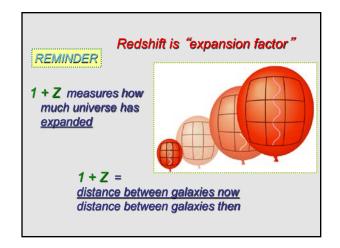




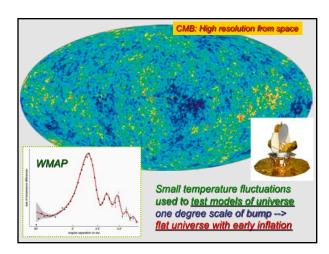


### Poll 1: looking back in time

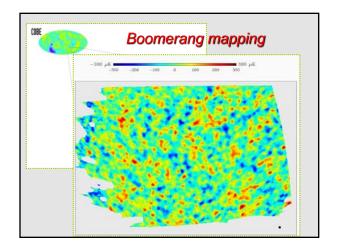
- If we can detect light from a quasar and decide that its emission line spectrum is at redshift Z = 4, how much bigger has the universe grown since that light left?
- A. 2 times bigger
- B. 3 times bigger
- C. 5 times bigger
- D. 16 times bigger

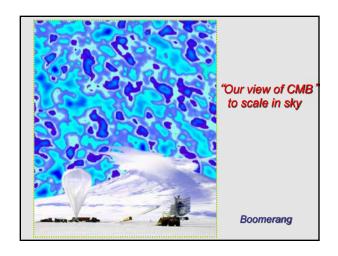


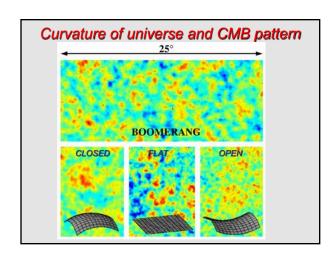
Questions or Comments

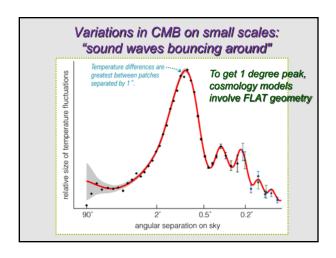


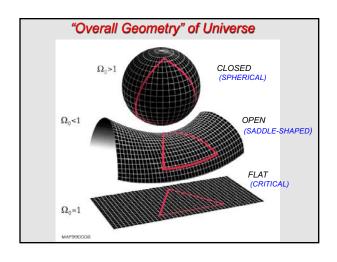












## What is the fate of the Universe?

- <u>Recollapse to gnaB giB ?</u>: crushing heat, destruction of all matter (Big Crunch) Rebirth ?
- <u>Eternal expansion</u>?: cold, galaxies dimming star formation slowing
- Everything winds up as a brown dwarf black dwarf, neutron star or black hole

Which is it?
Is there enough dark matter to recollapse the universe?

Baryonic matter: only few % of critical density

Dark matter: only about 25 % of what is needed

Looks like Universe is between the "coasting" and "critical" models

<u>Universe will expand forever</u> (or so it seems)

### **Questions or Comments**

