

Our Schedule

- Briefly revisit <u>collisions between galaxies:</u> "interacting galaxies" with bridges and tails
- Revisit "active galactic nuclei" (AGNs)
- Evidence for dark matter in galaxies
- Gravitational lensing: mainly by dark matter
- Finish reading Chap 22 Birth of Universe
- Start overview read Chap 23: Dark Matter, Dark Energy, Fate of Universe
- Focus on 23.2 Evidence for Dark Matter

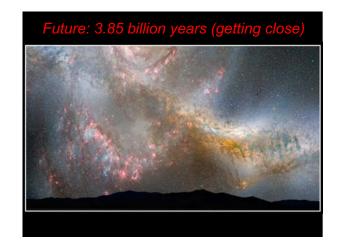


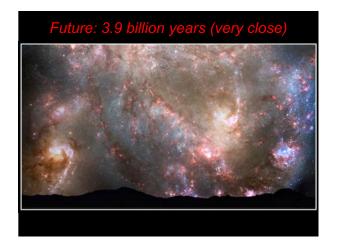








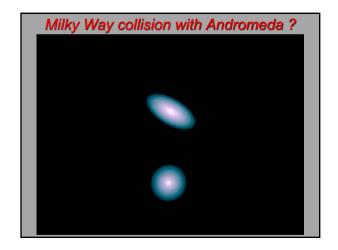


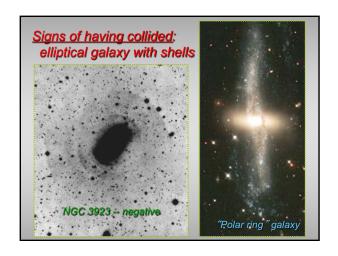


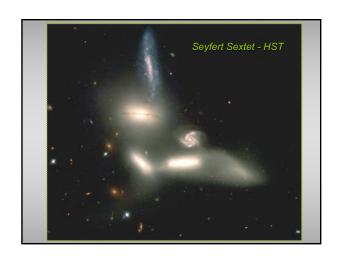


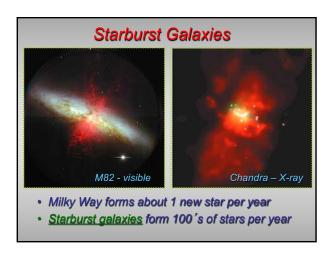


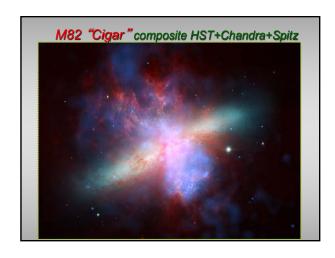


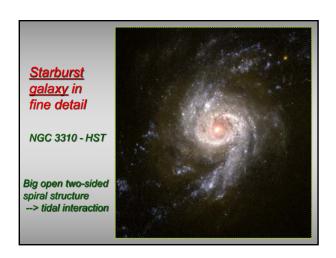


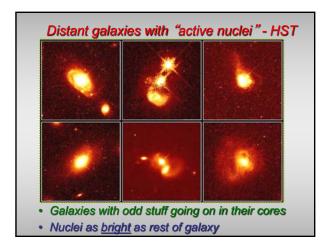


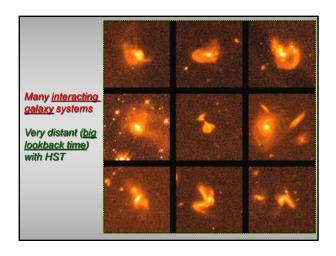






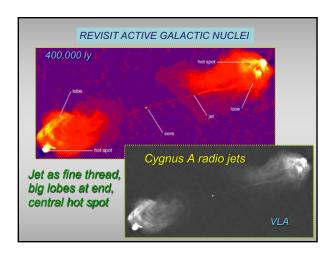


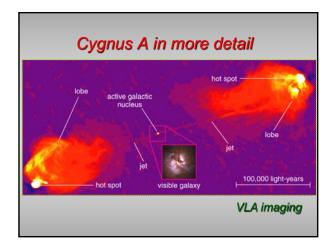


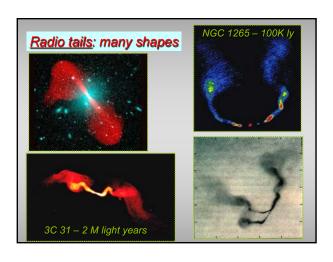


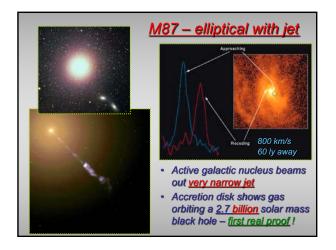
Messages from galaxy interactions

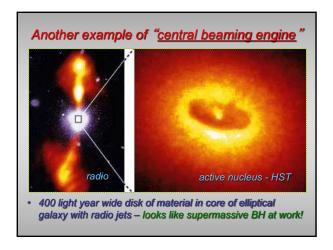
- 1. In <u>dense clusters</u>, galaxy collisions (grazing or even head-on) must have been common
- 2. With successive passages, spiral galaxies can tumble together to form a big elliptical
- 3. Vastly <u>increased star birth</u> from shocking the gas and dust (star burst galaxies)
- 4. Start <u>rapid feeding</u> of supermassive black hole lurking at center of most galaxies (quasars)

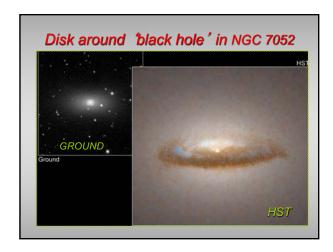


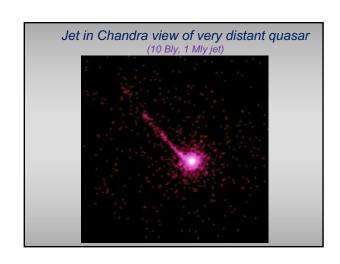












PROPERTIES OF "ACTIVE GALAXIES" RADIO GALAXIES, SEYFERT GALAXIES. RADIO GALAXIES , SEYFOLI GALANIE .

BL LACETAE OBJECTS , QUASARS "SOME HAVE MORE THAN OTHERS!" Typical properties of "active galaxies" 2. NON-THERMAL EMISSION

* EXCEST RADIATION IN UV, IR, RADIO, X-RAY

* IMPLIEUR SYNCHROTEON EMISTION FROM
BENITHISTIC ELECTRON'S STRUCKS IN MACAPIC
FIGURS synchrotron emission! 3. SMALL, COMPACT SIZE OF INTENSE EMISSION

O NUCLEUS VERY BRICHT COMPARED TO REST OF GALAXY source very 4. RAPIDLY VARYING EMISSION · Source May be A FEW LIGHT HOURS OR DAYS
IN SIZE small in size 5. EXPLOSIVE FEATURES

* JETUINE EXTENSIONS, FILAMENTS 6. GRAVITATIONAL DISTURBANCES Most quasars VELY HICH INTERNAL VELOCITIES DEPUGED FROM BROAD SPECIFIAL EMISSION LINES
 PECULIAL OPECAL APPERLANCE present when universe 7. LARGE REDSHIFTS

- IMPUNIS HIGH RECESTIONAL VELOCITIES,
VERY LARGE PISTANCES was young

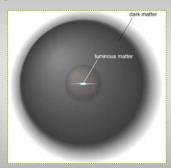
Clicker: galaxy collisions • Why are collisions between galaxies more likely than between stars within a galaxy? • A. Galaxies are much larger than stars • B. Galaxies travel through space much faster than stars • C. Relative to their sizes, galaxies are closer together than stars • D. Galaxies have higher redshifts than stars

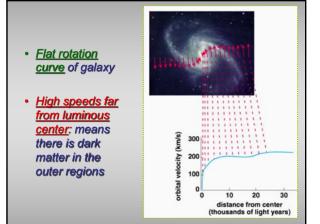
Now to Case for Dark Matter

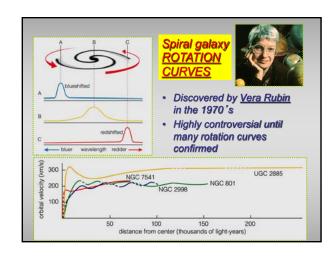
- ~ 80+% of <u>mass of universe</u> is dark matter (invisible, "missing" matter)
- Detectable ONLY via its gravitational forces on "lightful" matter (gas and stars)
- Note -- this dark matter is NOT the same as black holes, brown/black dwarfs, or dust

Individual galaxies show it

- <u>Rotation curves</u>: motions of stars in the galaxy
- Reveal that dark matter extends beyond visible part of the galaxy, mass is 10x stars and gas





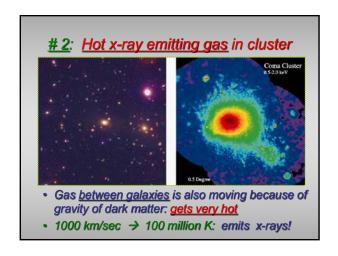


Galaxy Clusters: reveal dark matter in three ways

1: Galaxy velocities too large to be explained by gravity of visible galaxies

Expected ~100 km/sec for a typical cluster, found 1000 km/sec!

Discovered in 1930's by Fritz Zwicky (they didn't believe him, either)



Reading clicker - the boss galaxy

- Which of the following is <u>NOT</u> a feature of a central dominant (cD) galaxy in clusters?
 - lavios
- A. They are often spiral galaxies
- B. They are found in clusters of galaxies
- C. They often have multiple galactic nuclei near their centers
- D. They are thought to form by the merger of several smaller galaxies

