ASTR 1040: Stars & Galaxies

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Lecture 21 Tues 6 Nov 2018

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Today in Our Galaxy (Chap 19)

- Revisit *"inventory"* of our galaxy
- Why *spiral patterns are made in the disk* of galaxies, including our own
- Consider *rotation curve* of our galaxy, and the unseen mass *(dark matter)* that it implies
- Radio telescope *mapping with 21 cm H line*

- Overview read *Chap 20 "Galaxies and Foundations of Modern Cosmology"*
- *Observatory Reports* are coming due, before Fall Break (Thanksgiving)

REVISIT

Stars

Inventory of "stuff" making up our galaxy

Gas

+ "dark matter" as a halo

Dust

INVENTORY OF MILKY WAY

1. STARS — FEW HUNDRED BILLION, $\Sigma 10^{11} M_\odot$
   - BULGE — MEDIUM TO OLD, "METAL POOR"
   - DISK — YOUNG, "METAL RICH" — INCLUDES OPEN CLUSTERS, OB ASSOCIATIONS
   - HALO — OLDEST, METAL POOR — INCLUDES GLOBULAR CLUSTERS
2. GAS — 10% MASS OF STARS
   - MOSTLY IN DISK — INTERSTELLAR MEDIUM
     - A. VERY COLD GAS IN THIN SHEET — SITE OF STAR FORMATION (MOLECULAR CLOUDS)
     - B. WARM ATOMIC AND IONIZED H CLOUDS — EMISSION NEBULAE (BRIGHT NEBULAE)
     - C. HOT GAS — HEATED BY STELLAR WINDS, SUPERNOVAE
   - HALO — VERY HOT GAS BLOWING OUT OF GALAXY
3. DUST — 1% MASS OF GAS, 0.1% MASS OF STARS — MOSTLY IN CLOUDS IN DISK

Infrared light reveals stars whose visible light is blocked by gas clouds

Infrared

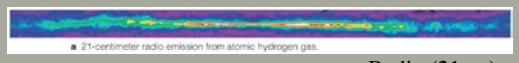
Visible

X-rays are observed from hot gas above and below the Milky Way's disk

We also see "point" sources – these are binaries

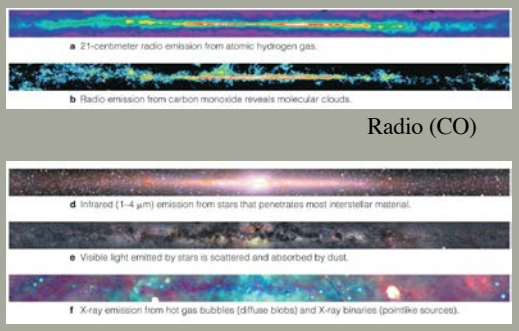
X-rays

21 cm radio waves emitted by atomic hydrogen show where gas has cooled and settled into disk



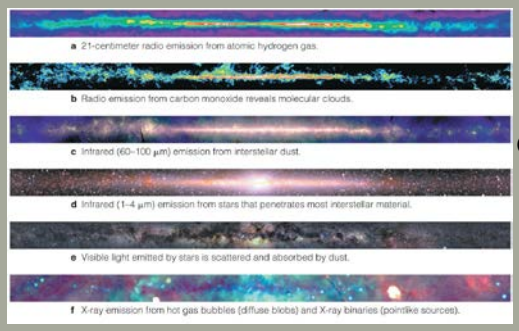
Radio (21cm)

Radio waves from carbon monoxide (CO) show locations of molecular clouds



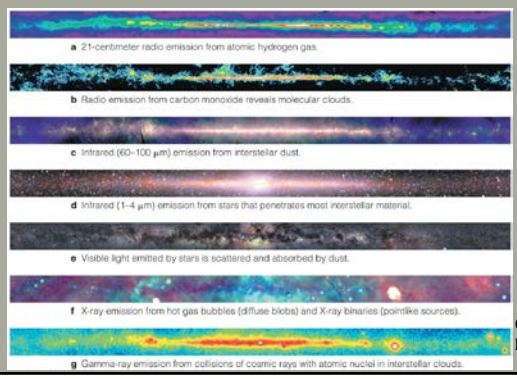
Radio (CO)

Long-wavelength infrared emission shows where young stars are heating the dust



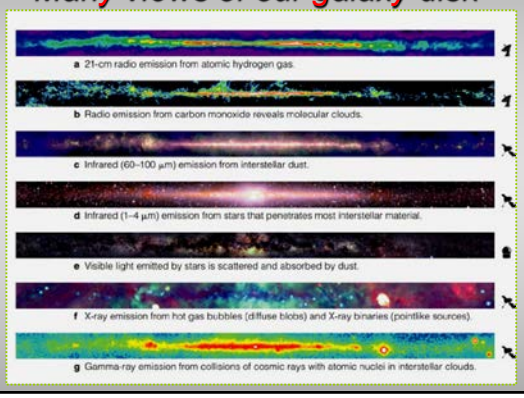
IR (dust)

Gamma rays show where cosmic rays from supernovae collide with atoms in gas clouds



Gamma-Rays

Many views of our galaxy disk

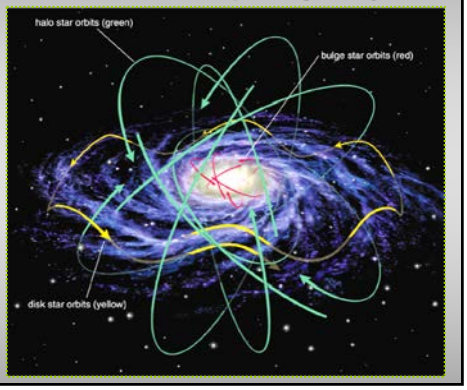


Motion of stars in spiral galaxy

REMIND

Halo & bulge: swarming in and out

Disk: circular + bit up/down



Clicker – stars and "heavy metals"

- *The ages of stars suggest that the bulge and halo of the Milky Way formed before many of the stars in the disk. Which would you expect to have more heavy metals (higher metallicity)?*

B.

- A. Halo and bulge stars
- B. Disk stars
- C. No difference

- **B.** Disk stars are continually forming out of gas that is more and more "polluted" by heavy metals.
- The OLD globular clusters in the halo were formed a long time ago before the galaxy was so polluted – they have very low "metallicities"

Why spiral arms?

"Density waves" – stars move in and out of denser regions

More like ripples in a pond than arms of a pinwheel

In dense regions, star formation is more intense, so "arms" are brighter

Push and pull of gravity in disk

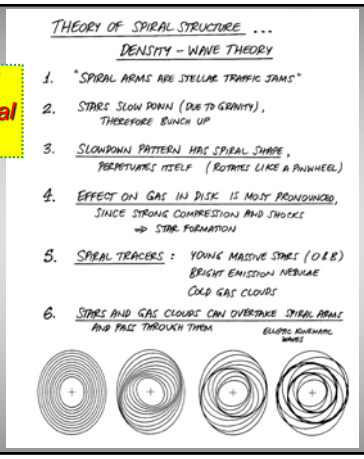
Gas/stars are pulled a little forward or backward toward the high density regions

Such clumping helps create a spiral pattern "traffic jam"

"Density wave" story – how spiral structure is built

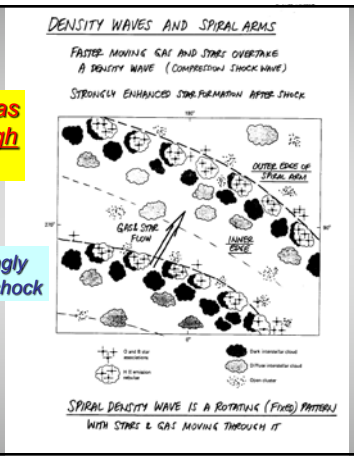
Gravitational instability of disks (gentle)

Alar Toomre "Q parameter" for disk instability



Stars and gas move through spiral wave

Star birth strongly enhanced by shock

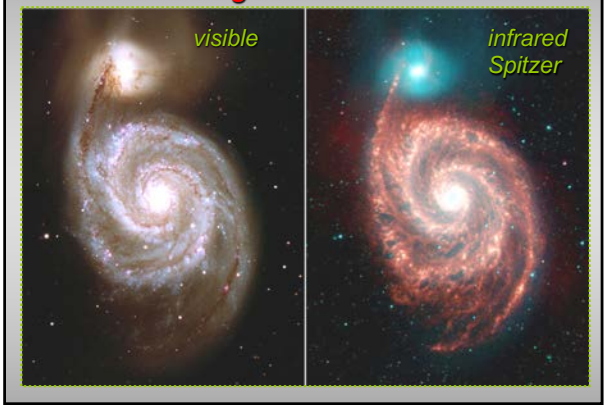


Crash/bang of star birth and recycling: rotating through the spiral arms in the disk

Bright O & B stars mark the spiral pattern

M51 Whirlpool Galaxy HST

Outlining starbirth in M51



Q: Why does "M51 – Whirlpool" have open spiral structure, and ours is pretty ragged?

A: Collision with companion galaxy enhanced its spiral wave

Clicker on deductions about Milky Way's stars

C.

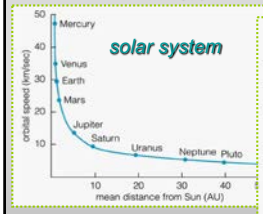
- Why are stars in the halo poor in the common elements carbon, nitrogen and oxygen?

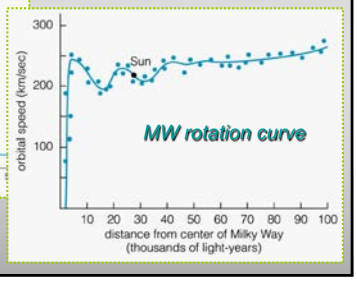
- A. Those elements have been used up in halo stars
- B. C, N and O are biological elements, and there is no life out there to make them
- C. The halo stars formed before these elements were made in abundance
- D. Making C, N and O requires massive stars, and these have been absent in the halo

Stars moving in circles: orbital velocity law

$$M_R = R \times V^2 / G$$

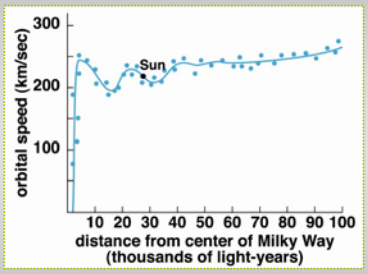
mass inward | radius | orbital velocity | …faster orbit, more mass





Reality for the Milky Way

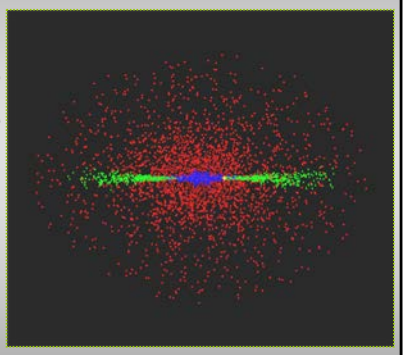
- Rotation curve is flat or even rising!
- Most of the mass of the galaxy is outside the solar circle!
- But few stars, little gas there…



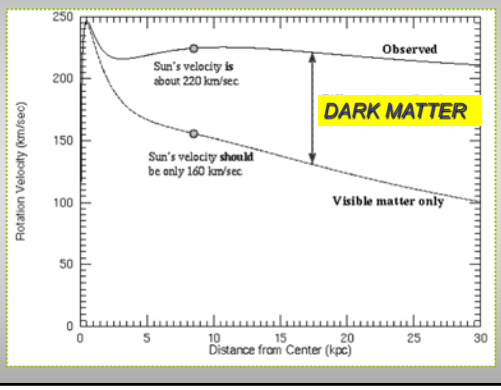
DARK MATTER ! probably in large halo -- outweights stars+gas by factor of about 10

Massive dark matter halo for MW

- Stars and gas are embedded in a much larger dark matter halo
- Don't know what dark matter is … probably not baryonic (usual protons, neutrons, electrons)

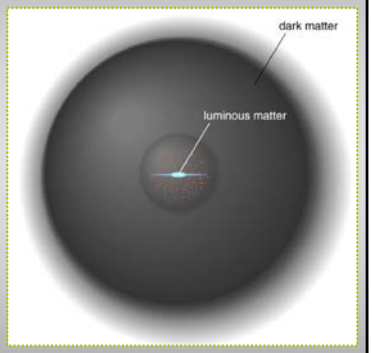


Role of dark matter on rotation profile



Dark matter halos for all galaxies

- Presence revealed by rotation curves (motions of stars in galaxy)
- Dark matter extends beyond visible part of the galaxy -- mass is ~10x stars and gas !
- Most likely subatomic particles, as yet unidentified (weakly interacting massive particles – WIMPs ?)

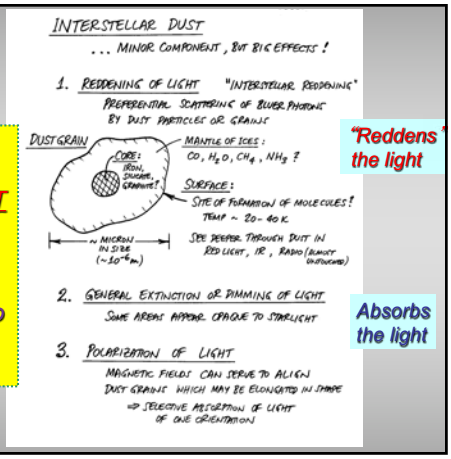


REVISIT

ISM:

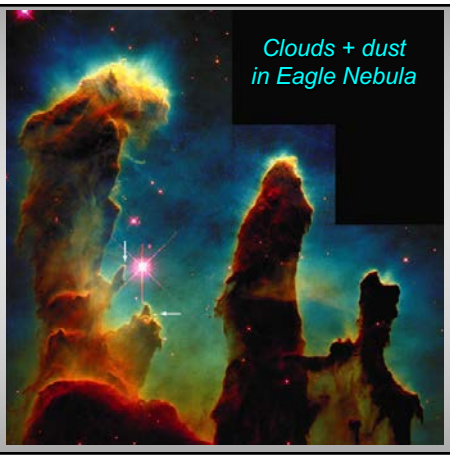
A little DUST goes a long way!

Confusing to distance & color



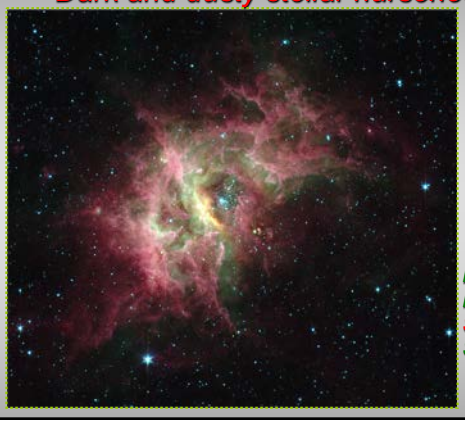
"Reddens" the light

Absorbs the light



Rosetta Nebula
O & B star carving + dust filaments

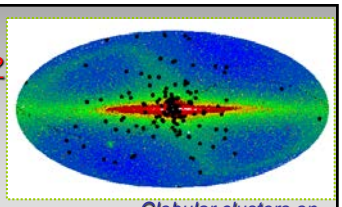
Dark and dusty stellar nurseries



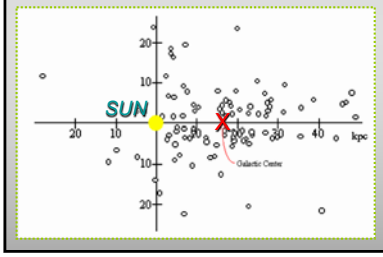
Nebula RCW 49
Stellar jewels
Spitzer IR

WHERE are we in MW?

Dust makes it hard to estimate distances (stars dimmed)



Globular clusters on COBE IR sky image



Shapley's (1920) distances to globular clusters: Sun at 18 kpc from center (twice too far!)

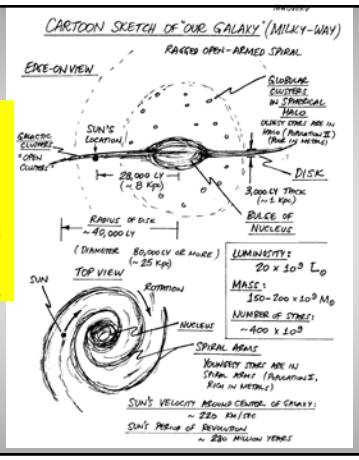
Large-scale structure in Milky Way

- We can observe the atomic hydrogen in interstellar gas in Milky Way with _____.

- A. space-based ultraviolet telescopes **D.**
- B. x-ray telescopes
- C. ground-based visible light telescopes
- D. 21 cm observations by radio telescopes

Q: Just HOW do we know that **MW has spiral structure?**

A: Radio astronomers could map it with **21-cm line of hydrogen**



21 cm emission (radio)

Sampling neutral hydrogen atoms (cold)

also how MRI does it in medical imaging!