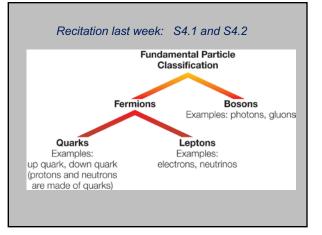


Topics for Today and Thur+

- Revisit solar magnetism and its cycles
- Use of supercomputers to simulate dynamics within the Sun
- Effects of solar magnetism on Earth
- What can we measure in other stars?
- How do we begin to classify other stars?
- Why <u>temperature and spectral lines</u> are closely linked in classifying stars O B A...M

Logistics

- Overview read Chap 15: Surveying the Stars
- <u>Review Session</u> Wed (tomorrow) 6-8pm here (G130) -- Max Weiner
- <u>Mid-Term Exam 1</u> Thurs in class (see rules in Review Set #1, still available)
- Homework #3 (+answers) returned today



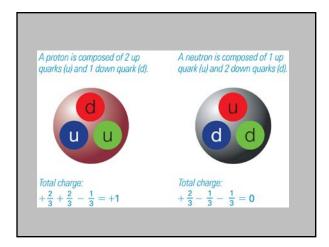
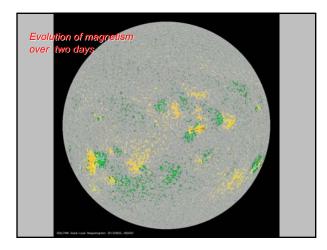
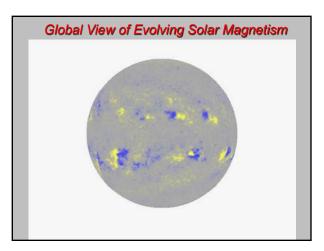
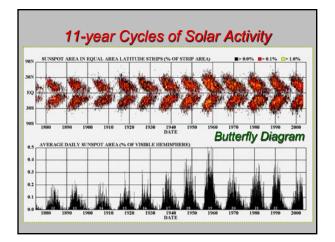
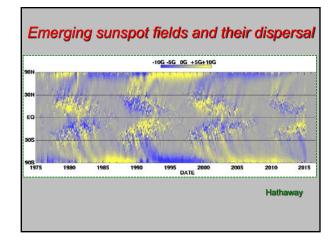


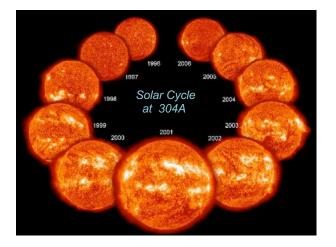
TABLE S4.1 Fundamental Fermions		
The Quarks	The Leptons	
Up	Electron	
Down	Electron neutrino	
Strange	Muon	
Charmed	Muon neutrino	
Тор	Tauon	
Bottom	Tau neutrino	

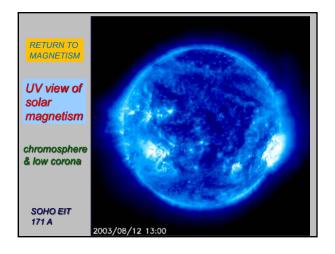


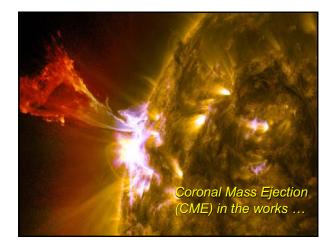


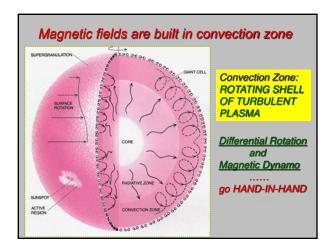


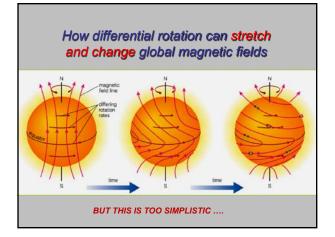






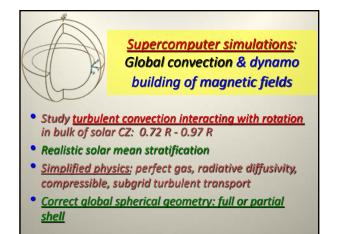


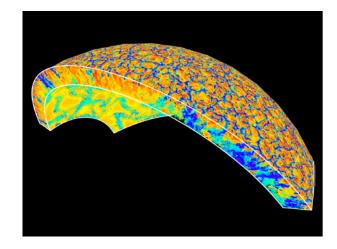


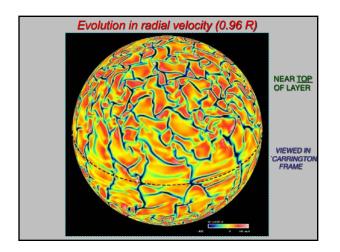


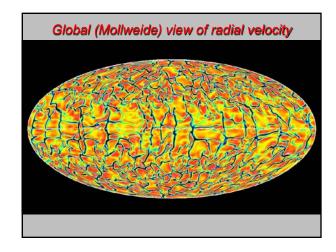
Reasoning Clicker Q B.

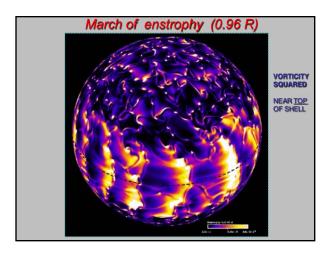
- If the Sun's core went out of balance and shrank a little, what would happen there?
- A. Density would decrease and fusion would slow down, releasing less energy
- **B.** Density and temperature would increase and fusion would <u>speed up</u>, releasing more energy
- C. The <u>whole</u> Sun would eventually shrink and thus core would come back into balance
- D. Not much would really change, so <u>nothing to</u> worry about

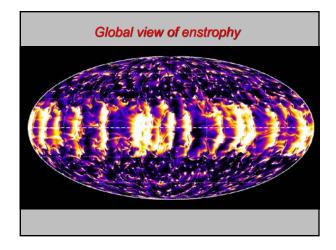


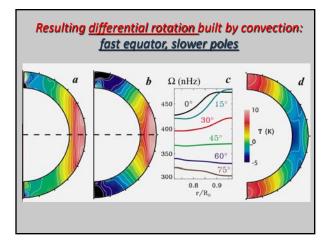


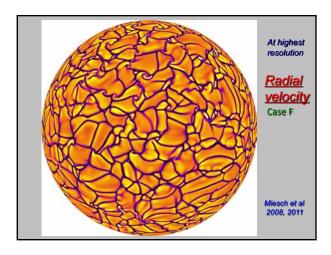


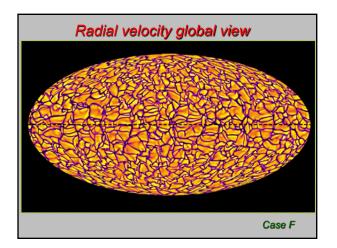


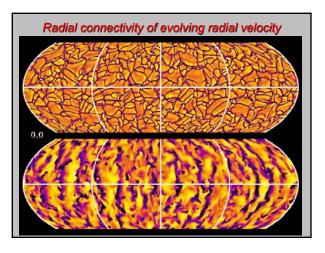


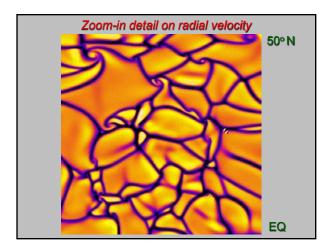


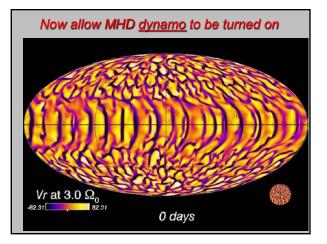


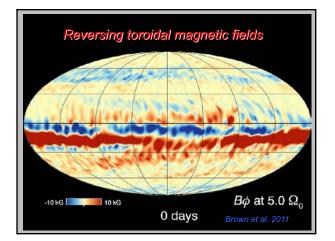


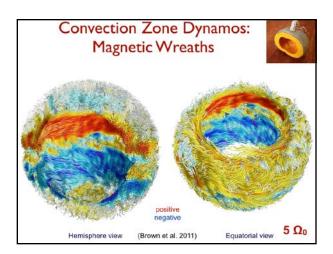


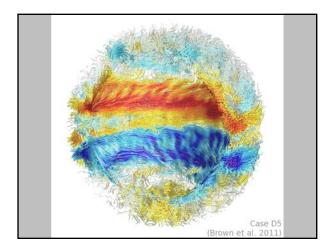


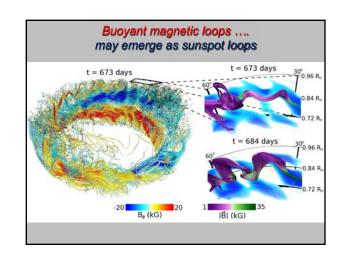




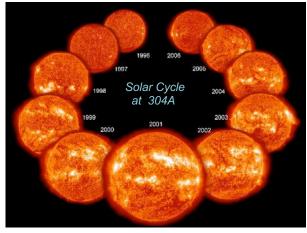


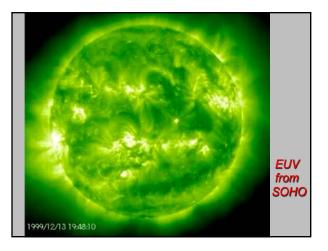


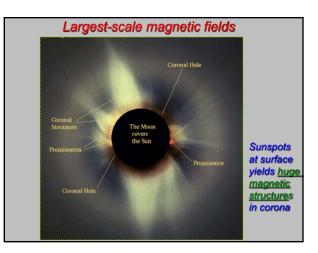


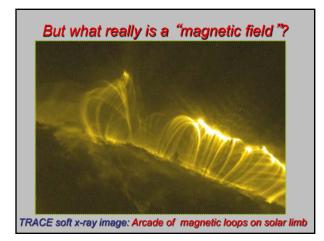












Reading Clicker Question Which is the most likely cause of the extreme heating in the chromosphere and corona?

- A. Energy deposited by magnetic fields
- B. Heat rising from the surface of the Sun
- C. Photons created at the photosphere interacting with the solar atmosphere
- D. Neutrino interactions with the solar wind
- E. Ionization of hydrogen just above the surface

